

FEniCS Course

Lecture 9: Incompressible Navier–Stokes equations

Contributors

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The incompressible Navier–Stokes equations

$$\begin{aligned} \dot{u} + u \cdot \nabla u - \nu \Delta u + \nabla p &= f && \text{in } \Omega \times (0, T] \\ \nabla \cdot u &= 0 && \text{in } \Omega \times (0, T] \\ u &= g_D && \text{on } \Gamma_D \times (0, T] \\ \nu \frac{\partial u}{\partial n} - pn &= g_N && \text{on } \Gamma_N \times (0, T] \\ u(\cdot, 0) &= u_0 && \text{in } \Omega \end{aligned}$$

- u is the fluid velocity and p is the pressure divided by the density ρ
- $\nu = \mu/\rho$ is the kinematic viscosity, μ dynamic viscosity
- f is a given body force per unit mass
- g_D is a given boundary velocity
- g_N is a given boundary function for the natural boundary condition
- u_0 is a given initial velocity

Variational problem

Multiply the momentum equation by a test function v and integrate by parts:

$$\begin{aligned} \int_{\Omega} (\dot{u} + u \cdot \nabla u) \cdot v \, dx + \nu \int_{\Omega} \nabla u : \nabla u \, dx - \int_{\Omega} p \nabla \cdot v \, dx \\ = \int_{\Omega} f \cdot v \, dx + \int_{\Gamma_N} g_N \cdot v \, ds \end{aligned}$$

Short-hand notation:

$$(\dot{u} + u \cdot \nabla u, v) + \nu(\nabla u, \nabla v) - (p, \nabla \cdot v) = (f, v) + (g_N, v)_{\Gamma_N}$$

Multiply the continuity equation by a test function q and sum up: find $(u, p) \in V$ such that

$$\begin{aligned} (\dot{u} + u \cdot \nabla u, v) + \nu(\nabla u, \nabla v) - (p, \nabla \cdot v) - (q, \nabla \cdot u) \\ = (f, v) + (g_N, v)_{\Gamma_N} \end{aligned}$$

for all $(v, q) \in \hat{V}$

Discrete mixed variational form of Navier–Stokes

Time-discretization leads to a *saddle-point* problem on each time step:

$$\begin{bmatrix} M + \Delta t A + \Delta t N(U) & \Delta t B \\ \Delta t B^\top & 0 \end{bmatrix} \begin{bmatrix} U \\ P \end{bmatrix} = \begin{bmatrix} b \\ 0 \end{bmatrix}$$

- Efficient solution of the saddle-point problem relies on the efficiency of special-purpose preconditioners (Uzawa iteration, Schur complement preconditioners, ...)
- We will use another approach (simpler and often more efficient)

The classical Chorin-Teman projection method

Step 1: Compute *tentative velocity* u^\star solving

$$\begin{aligned}\frac{u^\star - u^n}{\Delta t} - \nu \Delta u^\star + (u^\star \cdot \nabla) u^\star &= f^{n+1} && \text{in } \Omega \\ u^\star &= g_D && \text{on } \Omega_D \\ \frac{\partial u^\star}{\partial n} &= 0 && \text{on } \Omega_N\end{aligned}$$

Step 2: Compute a *corrected* velocity u^{n+1} and a *new* pressure p^{n+1} solving

$$\begin{aligned}\frac{u^{n+1} - u^\star}{\Delta t} + \nabla p^{n+1} &= 0 && \text{in } \Omega \\ \nabla \cdot u^{n+1} &= 0 && \text{in } \Omega \\ u^{n+1} \cdot n &= 0 && \text{on } \partial\Omega\end{aligned}$$

Computing the tentative velocity

In principle, the term $(u^* \cdot \nabla)u^{**}$ can be approximated in several ways

- Explicit: $u^* = u^{**} = u^n \Rightarrow$ diffusion-reaction equation
- Semi-implicit $u^* = u^n$ and $u^{**} = u^{n+1} \Rightarrow$ convection-diffusion-reaction equation
- Fully-implicit $u^* = u^{**} = u^{n+1}$ retaining the basic non-linearity in the Navier-Stokes equations

The natural outflow condition $\nu \partial_n u - pn = 0$ is artificially enforced by requiring

- $\partial_n u^\star = 0$ on $\partial\Omega_N$ in step 1
- $p^{n+1} = 0$ on $\partial\Omega_N$ in step 2

Solving the projection step

Applying $\nabla \cdot$ to $\frac{u^{n+1} - u^\star}{\Delta t} + \nabla p^{n+1} = 0$ and using requirement $\nabla \cdot u^{n+1} = 0$ yields

$$\Delta p^{n+1} = \frac{1}{\Delta t} \nabla \cdot u^\star \quad \text{in } \Omega$$

We already required

$$p = 0 \quad \text{on } \partial\Omega_N$$

Multiplying $\frac{u^{n+1} - u^\star}{\Delta t} + \nabla p^{n+1} = 0$ with n and restricting to $\partial\Omega_D$ gives

$$\frac{\partial p^{n+1}}{\partial n} = 0 \quad \text{on } \partial\Omega_D$$

Compute u^{n+1} by

$$u^{n+1} = u^\star - \Delta t \nabla p^{n+1}$$

including boundary conditions for u at $t = t^{n+1}$

Chorin-Teman projection method – Summary

- ❶ Compute tentative velocity u^\star by

$$\left(\frac{u^\star - u^n}{\Delta t}, v\right) + ((u^* \cdot \nabla)u^{**}, v) + \nu(\nabla u^\star, \nabla v) - (f, v) = 0$$

including boundary conditions for the velocity.

- ❷ Compute new pressure p^{n+1} by

$$(\nabla p^{n+1}, \nabla q) + \frac{1}{\Delta t}(\nabla \cdot u^\star, q) = 0$$

including boundary conditions for the pressure.

- ❸ Compute corrected velocity by

$$(u^{n+1} - u^\star, v) + \Delta t(\nabla p^{n+1}, v) = 0$$

including boundary conditions for the velocity.

Useful FEniCS tools (I)

Note `grad` vs. ∇ :

```
dot(grad(u), u)
dot(u, nabla_grad(u))
```

Solving linear systems:

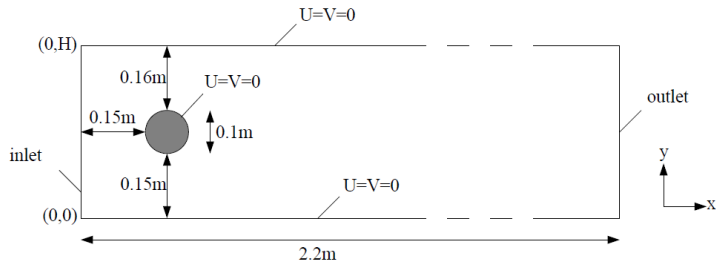
```
solve(A, x, b)
solve(A, x, b, "gmres", "ilu")
solve(A, x, b, "cg", "amg")
```

Defining a and L based on residual formulation:

```
F1 = ( (1/k)*inner(u - u0, v) + inner(grad(u0)*u0, v)
+ nu*inner(grad(u), grad(v)) - inner(f, v) ) * dx
a1 = lhs(F1)
L1 = rhs(F1)
```

The FEniCS challenge!

Implement a famous benchmark simulating a laminar flow around a cylinder. The geometry is described by

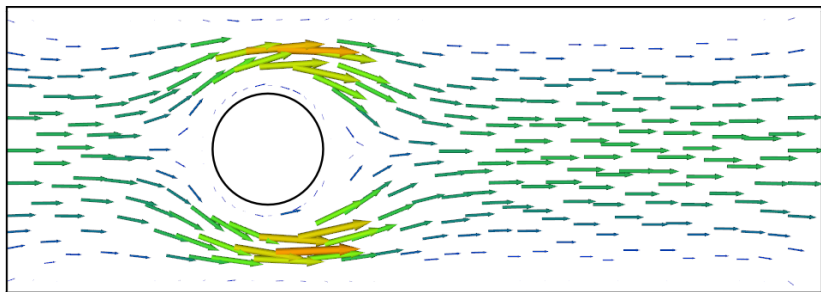


Set the kinematic viscosity $\nu = 0.001 \text{ m}^2/\text{s}$ and $\rho = 1.0 \text{ kg}/\text{m}^3$. A “do-nothing” boundary condition is assumed at the outlet. Defining $U_m = 1.5 \text{ m}/\text{s}$, the time-dependent inflow condition is given by

$$U = 4U_m y(H - y) \sin(\pi t/8)/H^2, \quad V = 0.$$

The FEniCS challenge!

The inflow boundary lies at $x = -0.2$ and the outflow boundary at $x = 2.0$. Compute the flow on the time interval $[0, 8]$ with time-step $dt = 0.001$. Test your implementation first for a larger time-step $dt = 0.01$ and the same channel problem but with the cylinder removed. If everything goes fine you should get something like



Happy coding!